

# Cosmics in the $\nu_e$ Appearance Search

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This talk includes work done by Christoph, Andrzej, and others.

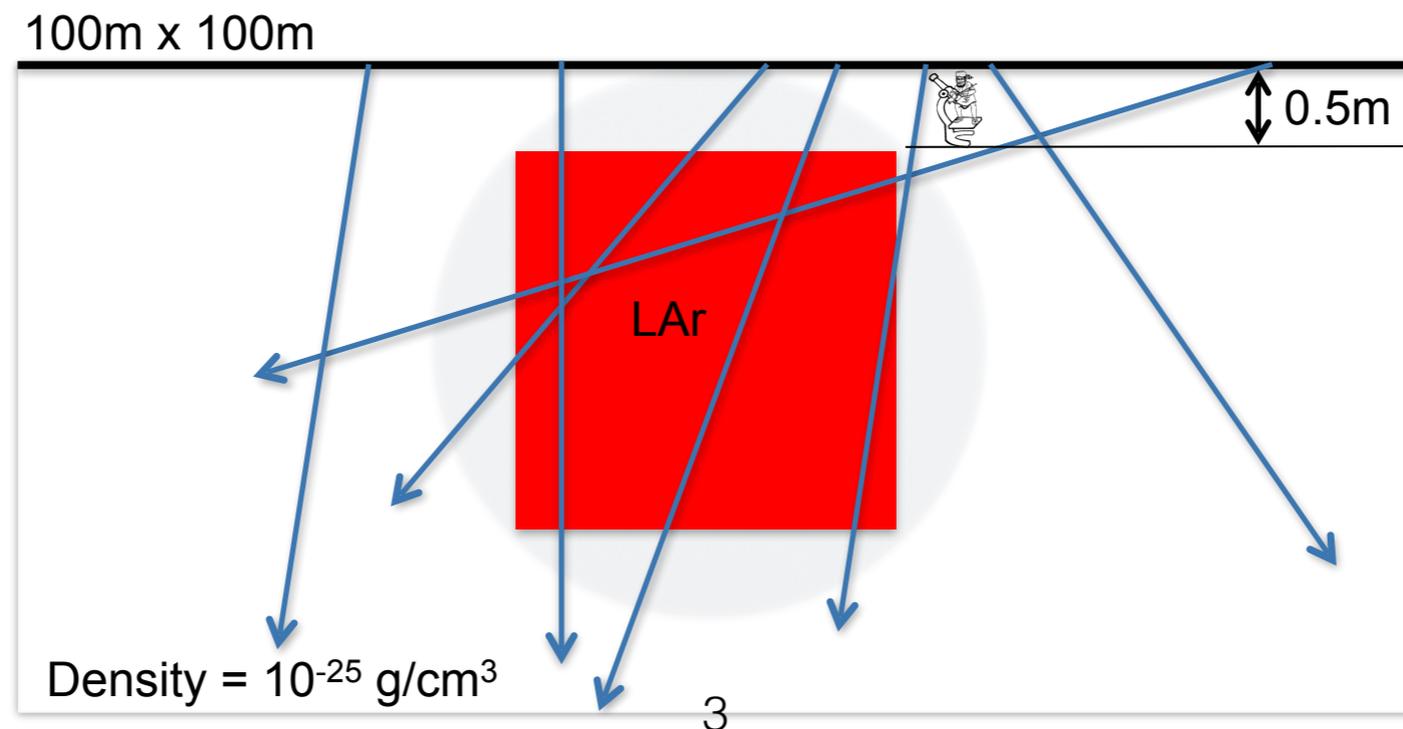


# What is the effect of Cosmic Backgrounds on MicroBooNE's ability to resolve an excess?

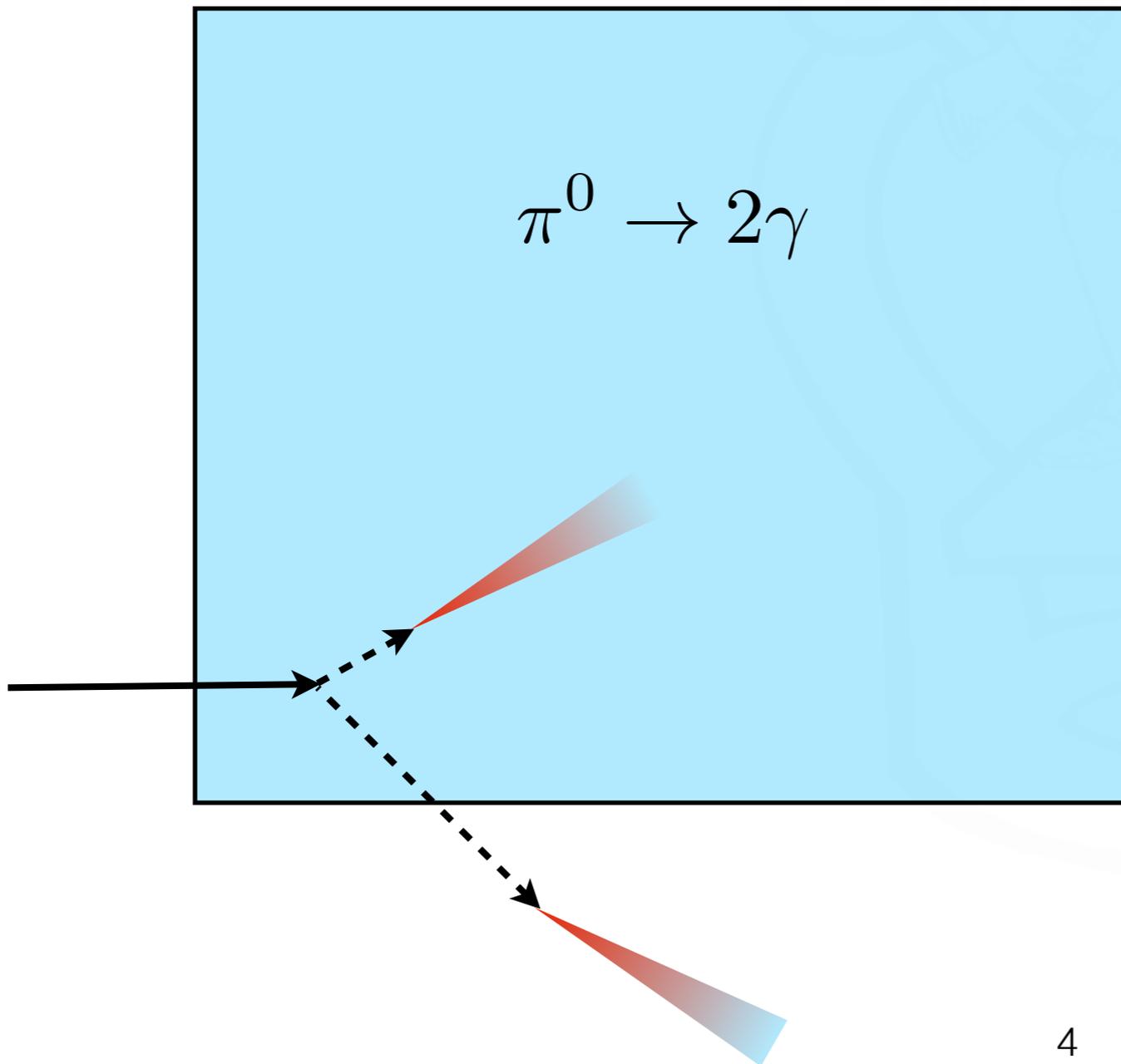
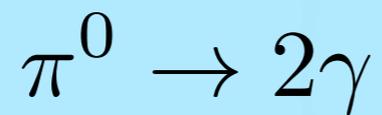
- Goal: combine the best simulation of cosmic events I can get with the beam Monte Carlo estimates to study the cosmic background in a  $\nu_e$  appearance search
- Examine the effect of a proposed, basic cut to mitigate the amount of cosmic backgrounds in the sample.
- Looking to get improved cosmic simulations from the Overburden Task force and incorporate them - the osc group can help that task force.

# Cosmic Sample

- Generated by Christoph using primary gammas from CRY and no other particles. Geometry is only the active volume of uBooNE.
- No gammas produced in the TPC itself
- **See MicroBooNE docdb [3438](#)**



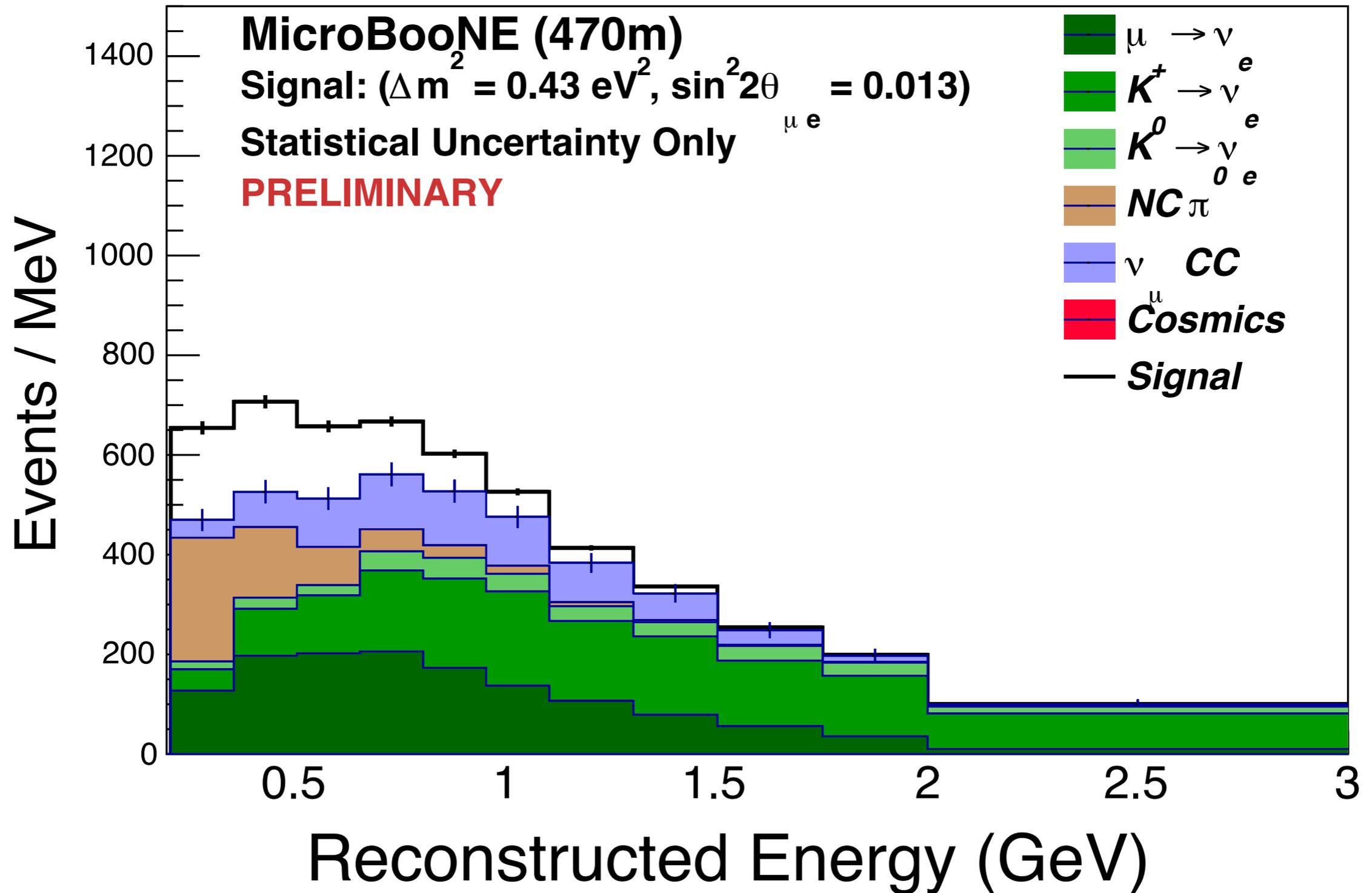
# $\nu_e$ Backgrounds



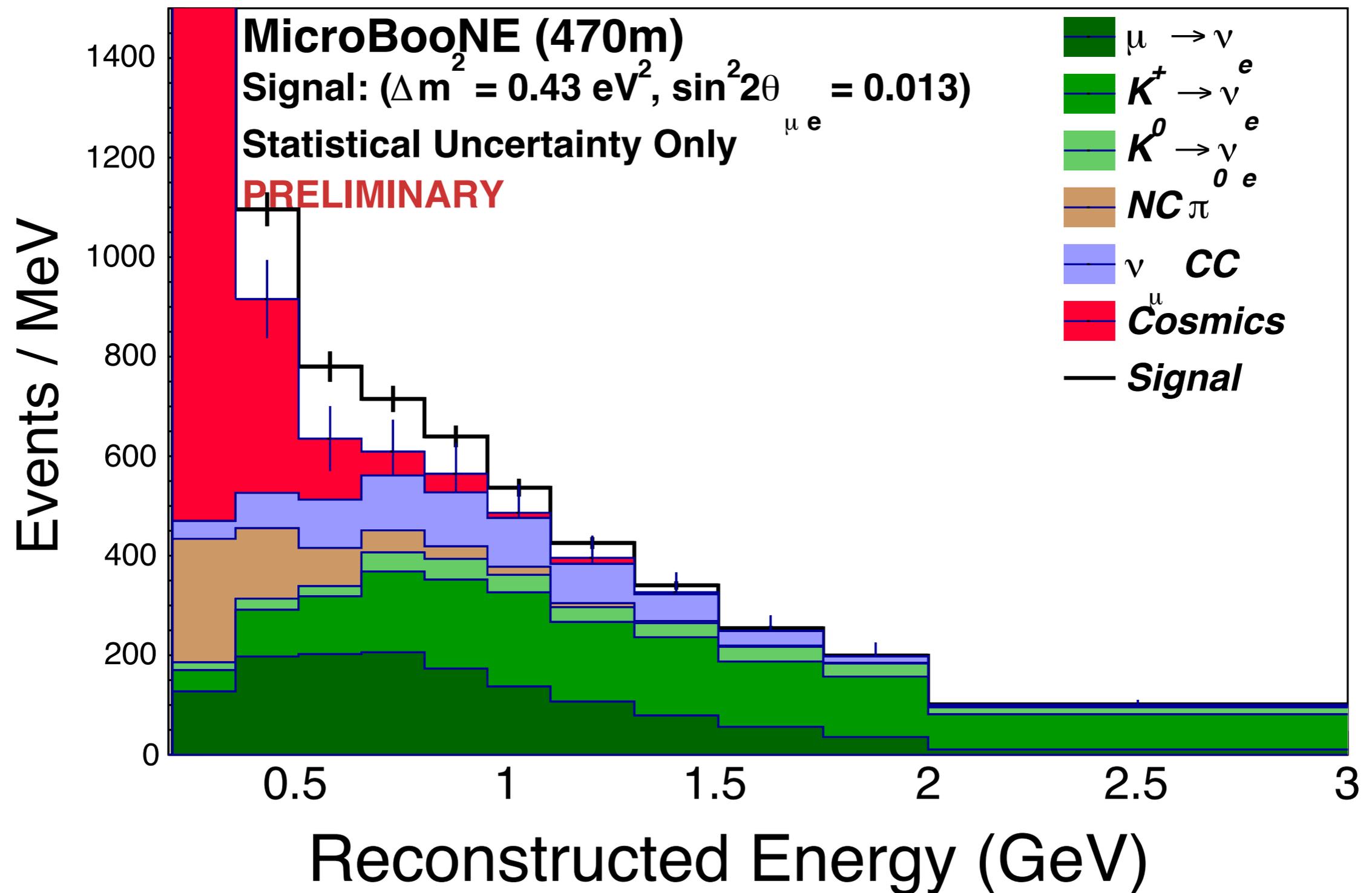
- Intrinsic Electron Neutrinos @ 80%
- $\nu_\mu$  CC MisID @ 0.1%
- Single Photons from  $\pi^0$  @ 6% (2 photons rejected)

Docdb 2963

# Nominal Distribution

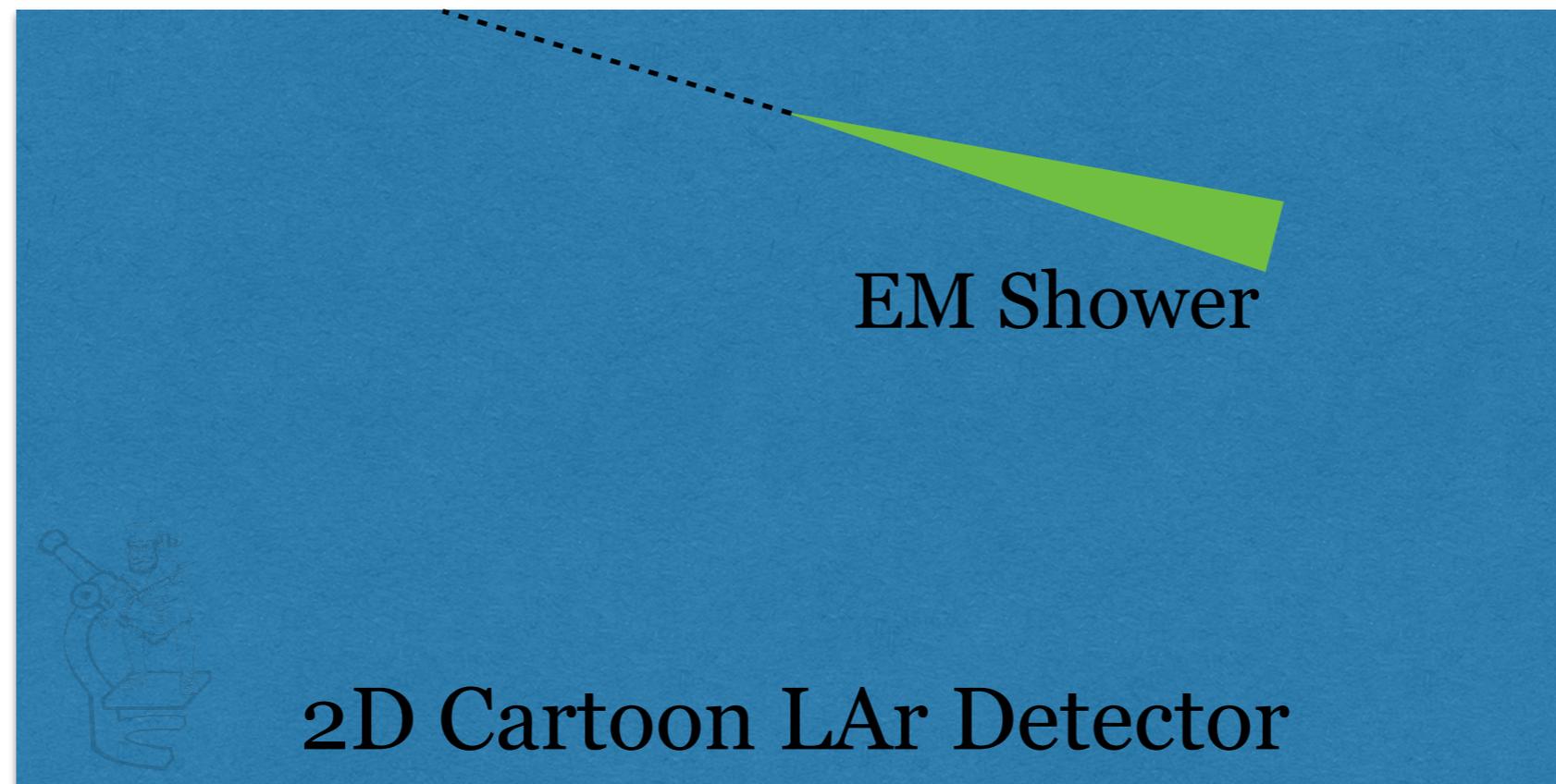


# Including all Cosmics



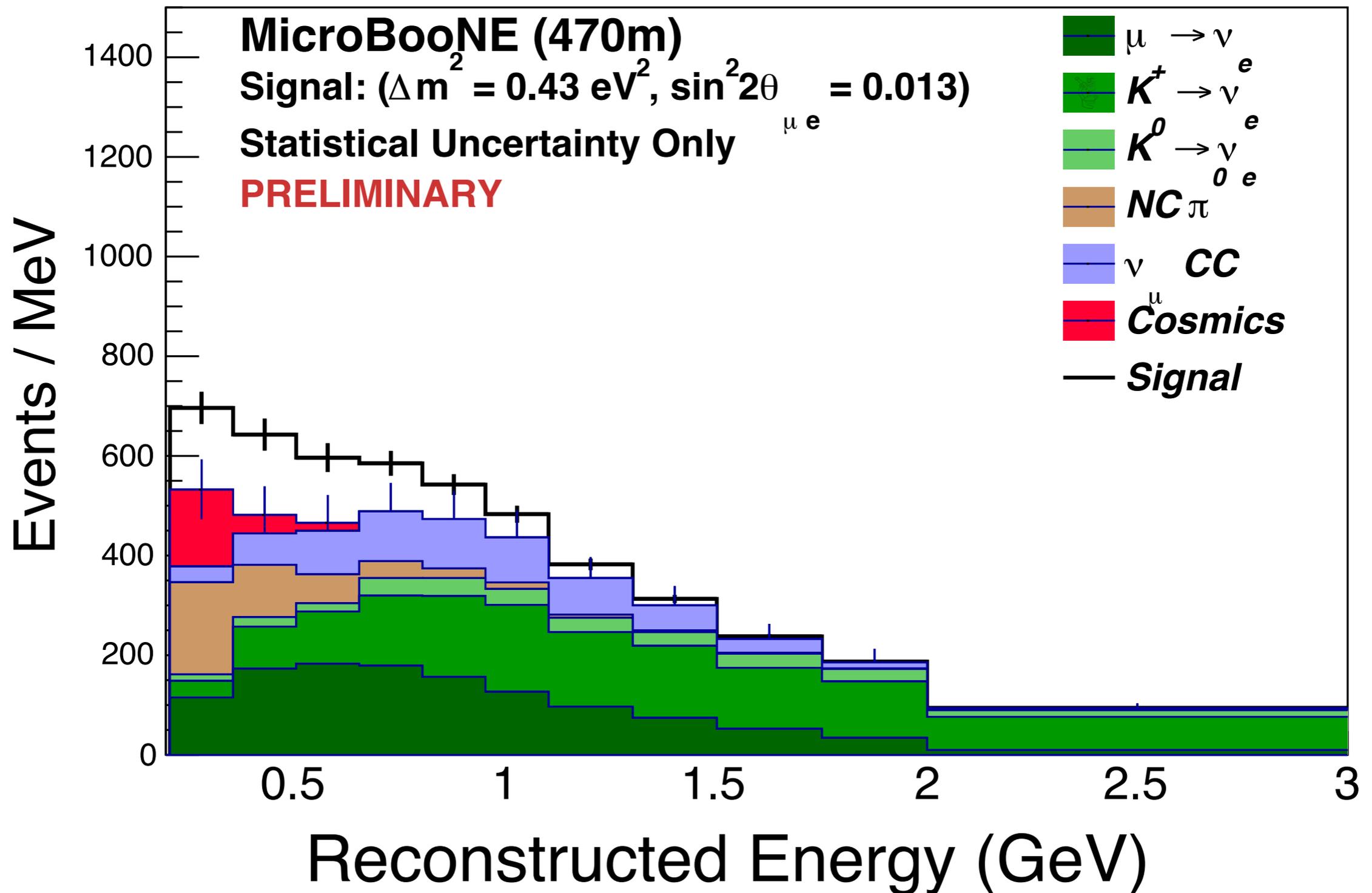
# Proposed Cut: “Backwards distance to Wall”

Distance, along axis, backwards  
towards detector wall

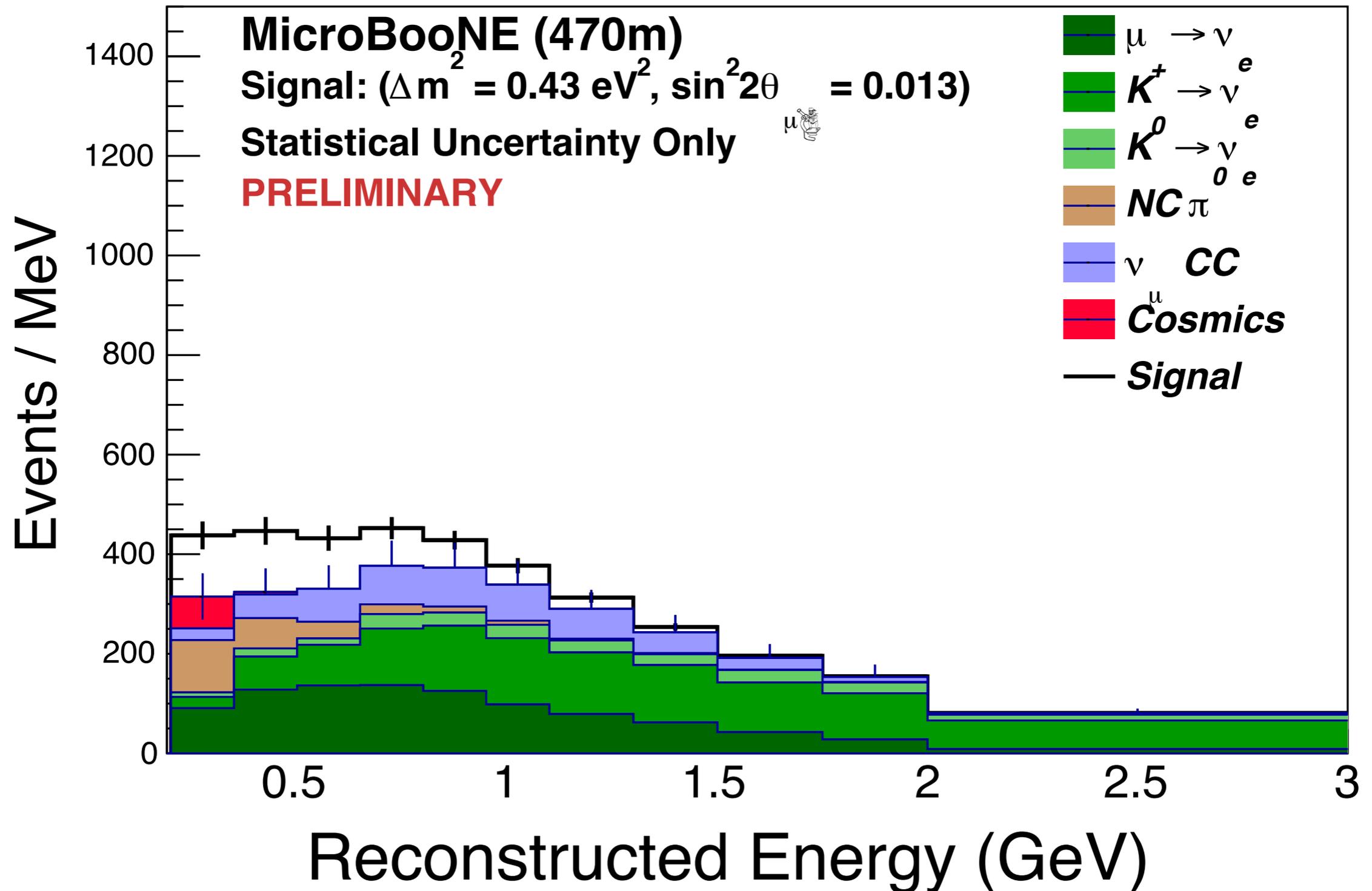


Assumption: Comptons from cosmic photons will occur  
closer to detector walls. Cut on the distance from the wall  
**along the axis of the shower.**

# 50cm Cut



# 100cm Cut



# Event rates for entire 200 MeV to 3GeV Region by Category

	No Cut	50 Cut	100 Cut	% Reduction @ 50cm	% Reduction @ 100cm
$\nu_e$ Intrinsic (total)	607	555	448	<b>9%</b>	<b>26%</b>
$\nu_e$ Intrinsic ( $\mu$ decay)	227	206	162	<b>9%</b>	<b>29%</b>
$\nu_e$ Intrinsic ( $K^+$ decay)	315	289	236	<b>8%</b>	<b>25%</b>
$\nu_e$ Intrinsic ( $K^0$ decay)	65	60	50	<b>8%</b>	<b>23%</b>
$\nu_\mu + e \rightarrow \nu_\mu + e$	6	6	5	<b>8%</b>	<b>14%</b>
$\pi^0$ Mis ID	87	66	38	<b>25%</b>	<b>57%</b>
$\nu_\mu$ Mis ID	118	108	86	<b>8%</b>	<b>27%</b>
Cosmics	352	31	10	<b>91%</b>	<b>97%</b>
Signal (3+1 Model)	116	105	81	<b>10%</b>	<b>30%</b>
<b>Signal / Sqrt(Bg.)</b>	<b>3.40</b>	<b>3.80</b>	<b>3.34</b>		

# Event Rates in Low Energy Region

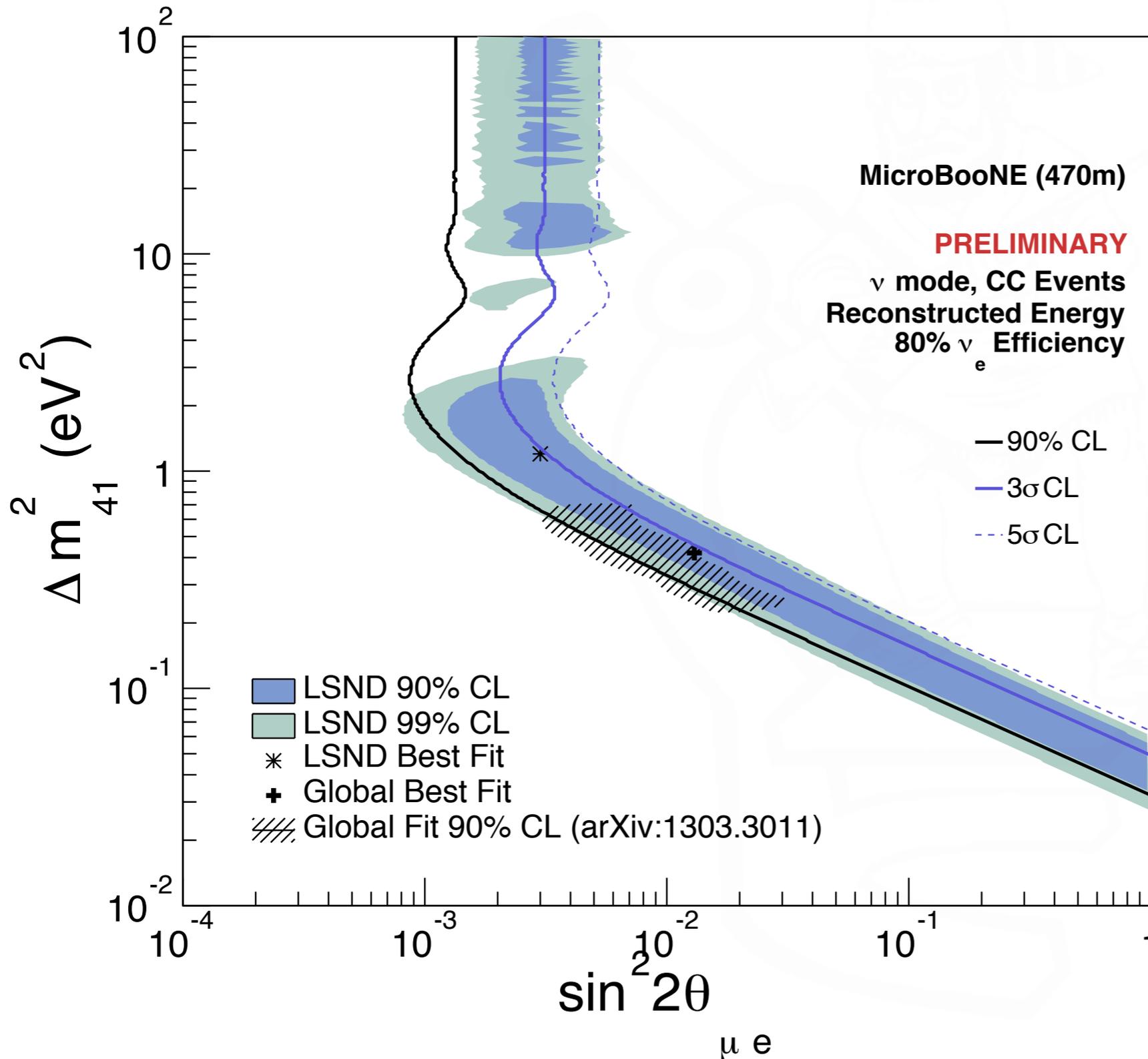
Range (MeV)	No Cut			50cm Cut			100cm Cut			% Red. @ 50cm	% Red. @ 100cm
	200-350	350-500	200-500	200-350	350-500	200-500	200-350	350-500	200-500		
$\nu_e$ Intrinsic (total)	28	47	<b>75</b>	24	41	<b>66</b>	18	32	<b>50</b>	<b>12%</b>	<b>33%</b>
$\nu_e$ Intrinsic ( $\mu$ decay)	19	30	<b>49</b>	17	26	<b>43</b>	14	19	<b>33</b>	<b>11%</b>	<b>33%</b>
$\nu_e$ Intrinsic ( $K^+$ decay)	6	14	<b>21</b>	5	13	<b>18</b>	3	10	<b>13</b>	<b>14%</b>	<b>35%</b>
$\nu_e$ Intrinsic ( $K^0$ decay)	2	3	<b>6</b>	2	3	<b>5</b>	1	2	<b>4</b>	<b>16%</b>	<b>33%</b>
$\nu_\mu + e \rightarrow \bar{\nu}_e + e$	2	1	<b>3</b>	1	1	<b>2</b>	1	1	<b>2</b>	<b>13%</b>	<b>23%</b>
$\pi^0$ Mis ID	37	21	<b>58</b>	28	16	<b>43</b>	16	9	<b>25</b>	<b>26%</b>	<b>57%</b>
$\nu_\mu$ Mis ID	5	11	<b>16</b>	5	9	<b>14</b>	4	7	<b>11</b>	<b>11%</b>	<b>34%</b>
Cosmics	257	58	<b>315</b>	23	6	<b>29</b>	10	1	<b>10</b>	<b>91%</b>	<b>97%</b>
Signal (3+1 Model)	26	26	<b>52</b>	23	23	<b>46</b>	17	17	<b>35</b>	<b>11%</b>	<b>33%</b>
Signal / Sqrt(Bg.)	1.44	2.20	<b>2.41</b>	2.59	2.69	<b>3.72</b>	2.49	2.46	<b>3.50</b>		

# Comments

- The neutral pions are reduced by more than beam intrinsic because the beam intrinsic events are evenly distributed throughout the active volume and the neutral pion misIDs are more likely to occur near the edge of the volume.
- This cut is very effective on these cosmics but does reduce intrinsic events. It could be optimized to maximize sensitivity to an excess by balancing misIDs with statistics.



# “Nominal Sensitivity”



To make comparisons between the different scenarios, it's useful to compute the sensitivity under the same assumptions for each scenario and compare the  $\chi^2$  at the same range of  $(\Delta m^2, \sin^2 2\theta)$ .

**Choose to use the 5 $\sigma$  contour in the nominal case as the reference set of points in  $(\Delta m^2, \sin^2 2\theta)$ .**

